

S1. Ans. (c)

$$g' = \frac{GM'}{R'^2} = \frac{GM}{10\left(\frac{R}{2}\right)^2}$$

$$= \frac{4}{10} \frac{GM}{R^2} = 0.4 \times 9.8$$

$$= 3.92 \text{ m s}^{-2}.$$

S2. Ans. (d)

Apply energy conservation,

$$U_i + K_i = U_f + K_f$$

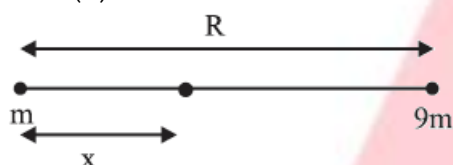
$$\Rightarrow -\frac{GMm}{R} + K_i = -\frac{GMm}{3R} + \frac{1}{2}mv^2$$

$$\Rightarrow -\frac{GMm}{R} + K_i$$

$$= -\frac{GMm}{3R} + \frac{1}{2} \times m \times \frac{GM}{3R}$$

$$\Rightarrow K_i = -\frac{1}{6} \frac{GMm}{R} + \frac{GMm}{R}$$

$$K_i = \frac{5}{6} \frac{GMm}{R}.$$

S3. Ans. (d)Let the gravitational field is zero at a distance x from the mass m .

$$\frac{Gm}{x^2} = \frac{G9m}{(R-x)^2}$$

$$\Rightarrow R-x = 3x \text{ or } x = \frac{R}{4}$$

Gravitational potential at $\frac{R}{4}$

$$= -\frac{Gm}{\frac{R}{4}} - \frac{G9m}{\frac{3R}{4}}$$

$$= -\frac{4Gm}{R} - \frac{12Gm}{R}$$

$$= -\frac{16Gm}{R}$$

S4. Ans. (c)

Time period of satellite

$$T = 2\pi \sqrt{\frac{R^3}{GM}}$$

$$= 2\pi \sqrt{\frac{R^3}{Gd \frac{4}{3}\pi R^3}}$$

$$\Rightarrow T = \sqrt{\frac{3\pi}{Gd}}$$

S5. Ans. (a)

$$I_g = \frac{F}{m}$$

$$= \frac{3}{60 \times 10^{-3}} = 50 \text{ N/kg}$$

S6. Ans. (a)

$$\text{Gravitational constant} = [M^{-1}L^3T^{-2}]$$

$$\text{Gravitational potential energy} = [ML^2T^{-2}]$$

$$\text{Gravitational potential} = [L^2T^{-2}]$$

$$\text{Gravitational intensity} = [LT^{-2}]$$

S7. Ans. (c)

$$\text{Hint: PE} + \text{KE} = mgs$$

At given point

$$\text{KE} = 3\text{PE}$$

$$\text{So, } 4\text{PE} = mgs$$

$$H = s/4$$

$$\text{KE} = \text{KE} = \frac{3mgs}{4} = \frac{1}{2}mv^2$$

$$V = \sqrt{\frac{3gs}{2}}$$

S8. Ans. (c)

$$\text{Hint: } V_e = \sqrt{\frac{2GM}{R}} = \sqrt{\frac{2G}{R}} \times \frac{4}{3}\pi R^3\rho$$

$$= \sqrt{\frac{8\pi G\rho}{3}} R$$

$$\Rightarrow V_e \propto R$$

$$\Rightarrow V_e/v = \frac{4R}{R} \Rightarrow V_e = 4v$$

S9. Ans. (c)

$$\text{Hint: } -\frac{GMn}{R} + \frac{1}{2}mk^2V_e^2 = -\frac{GMm}{R+r}$$

$$-\frac{GMn}{R} + \frac{1}{2}mk^2 \frac{2GMm}{R} = -\frac{GMn}{R+r}$$

$$-\frac{1}{R} + \frac{k^2}{R} = -\frac{1}{R+r}$$

$$\frac{1}{R+r} = \frac{1}{R} - \frac{k^2}{R}$$

$$\frac{1}{R+r} = \frac{1-k^2}{R}$$

$$r = \frac{Rk^2}{1-k^2}$$

S10. Ans. (a)

$$\text{Hint: } w_s = mg_s = 72 \text{ N}$$

$$w_h = mg_h = \frac{mg_s}{\left(1+\frac{h}{R}\right)^2} = \frac{72 \text{ N}}{\left(1+\frac{\frac{R}{2}}{R}\right)^2} = \frac{72}{9}$$

$$W_h = 32 \text{ N}$$

S11. Ans. (a)

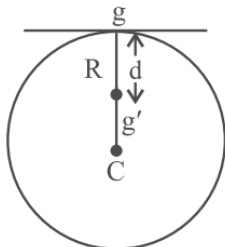
Hint: Inside the earth at depth d from the surface

$$g_{\text{eff}} = g \left(1 - \frac{d}{R}\right) \Rightarrow \frac{g}{n} = g \left(1 - \frac{d}{R}\right)$$

$$\Rightarrow d = \frac{(n-1)R}{n}$$

S12. Ans. (d)

Hint:



Acceleration due to gravity at a depth d from surface of earth

$$g' = g \left(1 - \frac{d}{R}\right) \dots\dots\dots(1)$$

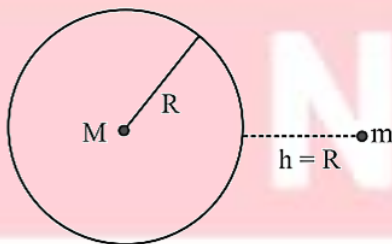
Here g = acceleration due to gravity at earth's surface Multiplying by mass ' m ' on both sides of a.

$$mg' = mg \left(1 - \frac{d}{R}\right); \left(\because d = \frac{R}{2}\right)$$

$$= 200 \left(1 - \frac{R}{2R}\right) = \frac{200}{2} = 100 \text{ N}$$

S13. Ans. (c)

Hint:



Initial potential energy

$$U_i = \frac{-GMm}{R}$$

Final potential energy at height $h = R$

$$U_f = \frac{-GMm}{(R+R)}$$

As work done = Change in PE

$$\therefore W = U_f - U_i$$

$$= \frac{GMm}{2R} = \frac{gR^2m}{2R} = \frac{mgR}{2} \quad (\because GM = gR^2)$$

S14. Ans. (b)

Hint: As angular momentum is conserved $mvr = \text{constant}$ so $v_A > v_B > v_C$

hence

$$K_A > K_B > K_C$$

S15. Ans. (d)

Hint: If universal constant becomes 10 times, then on earth surface

$$g = \frac{GM}{R^2} \quad G' = 10G$$

$$g' = 10 \frac{10GM}{R^2} \quad g' = 10g$$

So g on earth surface will change

S16. Ans. (c)

Hint: Acceleration due to gravity at height 1 km above earth surface

$$a_h = g \left[1 - \frac{2R}{R}\right]$$

$$= g \left[1 - \frac{2 \times 1}{R}\right]$$

$$= g \left[1 - \frac{2}{R}\right]$$

Acceleration due to gravity at depth ' d ' below earth surface

$$a_d = g \left[1 - \frac{d}{R}\right]$$

$$a_h = a_d$$

$$g \left[1 - \frac{2}{R}\right] = g \left[1 - \frac{d}{R}\right]$$

$$d = 2 \text{ km}$$

S17. Ans. (a)

Hint: Since two astronauts are floating in gravitational free space, the only force acting on the two astronauts is the gravitational pull of their masses

$$F = \frac{Gm_1m_2}{r^2}$$

which is attractive in nature. Hence they move towards each other.

S18. Ans. (b)

$$\text{Hint: } g_d = g \left[1 - \frac{d}{R}\right]$$

$$g_d = \frac{g}{4}$$

$$\frac{g}{4} = g \left[1 - \frac{d}{R}\right]$$

$$\frac{d}{R} = \frac{3}{4} \quad \dots\dots\dots(1)$$

$$g_h = g \left[\frac{R}{R+h} \right]^2$$

$$g_h = \frac{g}{4}$$

$$\frac{g}{4} = g \left[\frac{R}{R+h} \right]^2$$

$$\frac{h}{R} = 1 \dots\dots\dots(2)$$

From Eqs. (1) and (2), we get

$$\frac{d}{h} = \frac{3}{4} \qquad \frac{h}{d} = \frac{4}{3}$$

S19. Ans. (d)

Hint: Satellite in a orbit of radius R

$$TE = -\frac{GM_c m}{2R_E}$$

When satellite in a orbit of radii $3R_E$

$$TE_1 = -\frac{GM_c m}{2(3R_E)}$$

When satellite in a orbit of radii $9R_E$

$$TE_2 = -\frac{GM_c m}{2(9R_E)}$$

$$\Delta TE = TE_2 - TE_1$$

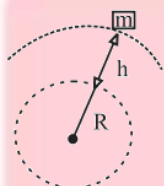
$$= -\frac{GM_c m}{18R_E} - \left[-\frac{GM_c m}{6R_E} \right]$$

$$= \frac{GM_c m}{R_E} \left[\frac{1}{6} - \frac{1}{18} \right]$$

$$= \frac{GM_c m}{9R_E}$$

S20. Ans. (d)

Hint:



$$PE = -\frac{GMm}{R+h}$$

$$KE = \frac{1}{2} m \left[\sqrt{\frac{GM}{R+h}} \right]^2$$

$$KE = \frac{GMm}{2(R+h)}$$

$$TE = \frac{-GMm}{R+h} + \frac{GMm}{2(R+h)}$$

$$= \frac{-GMm R^2}{2(R+h)R^2} \quad \left[\because g = \frac{GM}{R^2} \right]$$

$$= \frac{-g_0 m R^2}{2(R+h)}$$

S21. Ans. (d)

Hint: $g = \left(\frac{GM_e}{R_e^3} \right) r$ for $0 < r \leq R_e \Rightarrow g \propto r$

$g = \frac{GM_e}{r^2}$ for $r > R_e \Rightarrow g \propto \frac{1}{r^2}$

S22. Ans. (b)

Hint: g [in terms of mean density]

$$= \frac{4\pi G d R}{3} \Rightarrow V_e = \sqrt{2gR}$$

$$V_e = \sqrt{2 \times \frac{4}{3} \pi G d R \cdot R}$$

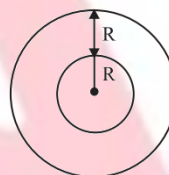
$$V_e = \sqrt{\frac{8}{3} \pi G d R^2}$$

$$V_p = \sqrt{\frac{8}{3} \pi G 2d (2R)^2} \because d_p = 2d, R_p = 2R$$

$$\frac{V_e}{V_p} = \frac{1}{\sqrt{8}} \Rightarrow \frac{1}{2\sqrt{2}}$$

S23. Ans. (a)

Hint:



Gravitational Potential

$$= \frac{-GM}{R+h} = -5.4 \times 10^7 \text{ J Kg}^{-2} \dots\dots\dots(1)$$

$$g = \frac{GM}{(R+h)^2} = 6 \text{ m/s}^2 \dots\dots\dots(2)$$

$$(1)/(2) = (R+h) = \frac{5.4 \times 10^7}{6} = 0.9 \times 10^7 \text{ m}$$

$$= 9000 \text{ km}$$

$$H = 9000 \text{ km} - 6400 \text{ km} = 2600 \text{ km}$$

S24. Ans. (a)

Hint: $T_p^2 = K r^3$ Given $\dots\dots\dots(1)$

$$T_{\text{planet}} = \frac{2\pi r}{V_p} \left[\because V_p = \sqrt{\frac{GM}{r}} \right]$$

$$= \frac{2\pi r \cdot r^{\frac{1}{2}}}{\sqrt{GM}}$$

$$T_p = \frac{2\pi r^{\frac{3}{2}}}{\sqrt{GM}} \dots\dots\dots(2)$$

Squaring Eq. (2), we have

$$T_p^2 = \frac{4\pi^2 r^3}{GM}$$

$$K r^3 = \frac{4\pi^2 r^3}{GM} \Rightarrow K = \frac{4\pi^2}{GM}$$

S25. Ans. (b)

Hint: For the satellite revolving around earth

$$v_0 = \sqrt{\frac{GM_e}{(R_e+h)}} \quad \sqrt{\frac{GM_e}{R_e\left(1+\frac{h}{R_e}\right)}} = \sqrt{\frac{gR_e}{1+\frac{h}{R_e}}}$$

Substituting the values

$$v_0 = \sqrt{60 \times 60^6} \text{ m/s}$$

$$v_0 = 7.76 \times 10^3 \text{ m/s} = 7.76 \text{ km/s}$$

S26. Ans. (a)

Hint:

Acceleration due to earth to the satellite is centripetal, hence directed towards centre. Angular momentum conservation holds good for comparable masses but

$$M_{\text{earth}} \gg M_{\text{satellite}}$$

S27. Ans. (a)

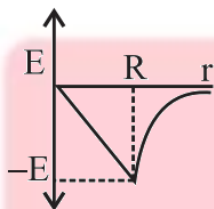
Hint: Gravitational field intensity

$$\vec{E} = \frac{-GM}{R^2}$$

$$\text{For a point inside the earth } E = \frac{-GMr}{R^3}$$

$$\text{For a point outside the earth } E = \frac{-GM}{R^2}$$

Where -ve sign indicates the attractive gravitational field



Accurate graph to show variation of E with r

S28. Ans. (a)

Hint: Escape velocity (v_e) = $\sqrt{\frac{2GM}{R}} = c$ = speed of light

$$\Rightarrow R = \frac{2GM}{c^2} = \frac{2 \times 6.6 \times 10^{-11} \times 5.98 \times 10^{24}}{(3 \times 10^8)^2} \text{ m}$$
$$= 10^{-2} \text{ m}$$

S29. Ans. (a)

$$\text{Hint: } V = -G(2) \left[\frac{1}{1} + \frac{1}{2} + \frac{1}{4} + \frac{1}{8} + \dots \right]$$

Because it forms geometric progression

$$S_{GP} = \frac{r^n - 1}{r - 1}; \quad r = \frac{1}{2}$$

$$\frac{\left(\frac{1}{2}\right)^n - 1}{-\frac{1}{2}} = \frac{-1}{-\frac{1}{2}} = 2$$

$$\therefore V = -2G \times S_{GP}$$

$$= -2G \times 2 = -4G$$

S30. Ans. (c)

$$\text{Hint: Change in P.E.} = -\frac{GMm}{3R} - \left(-\frac{GMm}{R}\right)$$

$$= \frac{2}{3} \frac{GMm}{R} = \frac{2}{3} mgR$$